# A THEORY OF LINEARIZED EXPLOSION PROBLEMS InCluding back pressure 

## (K TEORII LINEARIZIROVANNYKH ZADACH O VZRYVE C UCHETOM PROTIVODAVLENIIA)

PMM Vol.23, No.4, 1959, pp. 749.759<br>V.P. KOROBEINIKOV and E.V. RIAZANOV<br>(Moscow)<br>(Received 7 February 1959)

The statement and complete analytical solution of the one-dimensional self-similar problem of a strong point explosion in a gas is given by Sedov [1] (see also [2]). Here it was assumed that the initial density of the gas $\rho$, either is constant or depends on the geometrical coordinate $r$ according to the following law:

$$
\begin{equation*}
\mathrm{p}_{1}(r)=A r^{-\omega}, \quad \omega, \quad A-\text { const }, \quad A>0 \tag{0.1}
\end{equation*}
$$

In addition, it was considered that the initial pressure $p_{1}$ of the undisturbed gas can be neglected in comparison with the pressure $p_{2}$ at the shock wave front. It is known that the latter assumption is correct only for the initial stage of the explosion development, i.e. for a brief time interval. As the shock wave propagates the influence of the initial pressure becomes essential. Therefore, in the initial conditions of the problem there appears an additional dimensional parameter $p_{1}$, by virtue of which the problem ceases to be self-similar and all non-dimensional characteristics of the flow will now depend not on one but on two variables. The non self-similar problem can be solved by numerical integration of a system of nonlinear equations, and this has been done for the case of constant initial density [3] with the help of high-speed computing machines.

However, a simpler means of solving the problem for values of ( $p_{2}-p_{1}$ )/ $p_{1}<1$ can be proposed, which is based on a linearization of the basic equations about the self-similar solution. For the problem of an explosion in a perfect gas with constant density and adiabatic exponent $\gamma=1.4$ linearized solutions have been obtained previously [1, 4,5]. In those papers the linearized non-dimensional parameter is $q$, which characterizes the intensity of the shock wave and is equal to the ratio of the square of the velocity of sound in the undisturbed gas to the square of the velocity of the shock wave:

$$
\begin{equation*}
q=a_{1}^{2} c^{-2}=\gamma p_{1} p_{1}{ }^{-1} c^{-2} \tag{0.2}
\end{equation*}
$$

In the present paper a solution of the linearized problem of an explosion taking into account back pressure in a medium with varying initial density, which is determined by formula ( 0.1 ), will be considered. Using the method of linearization on the variable parameter $q$, equations which describe one-dimensional motions similar to the self-similar cases are derived. A first integral of the obtained system of linearized equations is found and an exact analytical solution of the problem is given for

$$
\begin{equation*}
\omega=\omega_{1}=\frac{3 v-2+\gamma(2-v)}{\gamma+1}, \tag{0.3}
\end{equation*}
$$

where $\nu=3$ for spherical waves, $\nu=2$ for cylindrical waves, and $\nu=1$ for plane waves.

1. Statement of the problem and the basic equations. We shall take the initial system of equations of gas dynamics, which describe the one-dimensional adiabatic disturbed motions of a perfect gas behind the shock wave front, in the form

$$
\begin{align*}
& \frac{\partial v}{\partial t}+v \frac{\partial v}{\partial r}+\frac{1}{\rho} \frac{\partial p}{\partial r}=0 \\
& \frac{\partial \rho}{\partial t}+v \frac{\partial \rho}{\partial r}+\rho\left(\frac{\partial v}{\partial r}+\frac{v-1}{r} v\right)=0  \tag{1.1}\\
& \frac{\partial p}{\partial t}+v \frac{\partial p}{\partial r}+\tau p\left(\frac{\partial v}{\partial r}+\frac{v-1}{r} v\right)=0
\end{align*}
$$

For the solution of the problem of a point explosion it is required to find a solution of system (1.1) with the boundary conditions at the shock wave front:

$$
\begin{equation*}
v_{2}=\frac{2}{\gamma+1}(1-q) c, \quad \rho_{2}=\frac{\gamma+1}{\gamma-1+2 q} \rho_{1}, \quad p_{2}=\frac{2 \gamma-(\gamma-1) q}{(\gamma+1) q} p_{1} \tag{1.2}
\end{equation*}
$$

Quantities immediately behind the shock wave front are denoted by the index 2. Let $r_{2}(t)$ be the radius of the shock wave. Then

$$
v_{2}=v\left(r_{2}, t\right), \quad \rho_{2}=\rho\left(r_{2}, t\right), \quad p_{2}=p\left(r_{2}, t\right), \quad c=d r_{2} / d t
$$

The dependence of $v_{2}(t), \rho_{2}(t), p_{2}(t), r_{2}(t)$ on time is unknown beforehand, their determination being equivalent to the determination of the dependence of $q(t)$. In addition to conditions (1.2), we also have the boundary condition for the velocity at the center of symmetry

$$
\begin{equation*}
v(0, t)=0 \tag{1.3}
\end{equation*}
$$

At time $t=0$ a finite energy $E_{0}$ is released at the center of symmetry and the initial conditions

$$
r(r, 0)=0, \quad \rho(r, 0)=\rho_{1}(r)=A r^{-\omega}, \quad p(r, 0)=p_{1}=\mathrm{const}, \quad r_{2}(0)=0(1.4)
$$

are given.
From the system which defines the parameters of this problem $A, p_{1}$, $E_{0}, \gamma, \omega, r, t$ it follows that the required non-dimensional functions

$$
\begin{equation*}
f=v / c, \quad g=p ; p_{2}, \quad h=p_{i}^{\prime} p_{2} \tag{1.5}
\end{equation*}
$$

will depend on two non-dimensional variables, for which we take

$$
\lambda=r / r_{2}, \quad q=a_{1}^{2} / c^{2}
$$

and on the constant parameters $\gamma$ and $\omega$.
Passing to the introduced non-dimensional variables and taking into account that

$$
\begin{gathered}
\frac{\partial}{\partial r}=\frac{1}{r_{2}} \frac{\partial}{\partial \lambda}, \quad \frac{\partial}{\partial t}=\frac{c}{r_{2}}\left(r_{2} \frac{d q}{d r_{2}} \frac{\partial}{\partial q}-\lambda \frac{\partial}{\partial \lambda}\right) \\
\frac{2}{c^{2}} \frac{d c}{d t}=\frac{2}{c} \frac{d c}{d r_{2}}=\frac{d \ln c^{2}}{d r_{2}}=\left(\omega-\frac{r_{2}}{q} \frac{d q}{d r_{2}}\right) r_{2} \\
\frac{1}{c} \frac{d \ln p_{2}}{d t}=\frac{d \ln p_{2}}{d r_{2}}=-\frac{1}{r_{2}}\left(\omega+\frac{2}{\gamma-1+2 q} r_{2} \frac{d q}{d r_{2}}\right) \\
\left.\frac{1}{c} \frac{d \ln p_{2}}{d t}=\frac{d \ln p_{2}}{d r_{2}}=-1 \overline{[2 \gamma}-(\gamma-1) q\right] q d r_{2}
\end{gathered}
$$

system (1.1) can be transformed into the form

$$
\begin{align*}
& (f-\lambda) \frac{d f}{d \lambda}+\frac{(\gamma-1+2 q)[2 \gamma-(\gamma-1) q]}{\gamma(\gamma+1)^{2}} \frac{1}{g} \frac{\partial h}{\partial \lambda} \div\left(\frac{\partial f}{\partial q}-\frac{f}{2 q}\right) r_{2} \frac{d q}{d r_{2}}+\frac{\omega}{2} f=0 \\
& (f-\lambda) \frac{\partial \ln g}{\partial \lambda}+\frac{\partial f}{\partial \lambda}+\frac{v-1}{\lambda} f+\left(\frac{\partial \ln g}{\partial q}-\frac{2}{\gamma-1-2 q}\right) r_{2} \frac{d q}{d r_{2}}-\omega=0  \tag{1.6}\\
& (j-\lambda) \frac{\partial \ln h}{\partial \lambda}+\gamma\left(\frac{\partial f}{\partial \lambda}+\frac{v-1}{\lambda} f\right)+\left(\frac{\partial \ln h}{\partial q}-\frac{2 \gamma}{[2 \gamma-(\gamma-1) q] q}\right) r_{2} \frac{d q}{d r_{2}}=0
\end{align*}
$$

We shall define the non-dimensional radius $R_{2}$ of the shock wave by the formula

$$
R_{2}=r_{2} r^{\circ} \quad\left(r^{\circ}=\left(E_{0} p_{1}\right)^{\frac{1}{v}}\right)
$$

where $r^{0}$ is a characteristic dynamic length.
In order to obtain the complete solution of the problem in the adopted variables, it is necessary to determine $f(\lambda, q), g(\lambda, q), h(\lambda, q)$ and also of $R_{2}(q)$. For this it is necessary to find a solution of system (1.6) in the $\lambda, q$ plane inside the square $0 \leqslant \lambda \leqslant 1$ and $0 \leqslant q \leqslant 1$ which satisfies the following boundary conditions:
at the shock

$$
\begin{equation*}
f(1, q)=g(1, q)=h(1, q)=1 \quad \text { for } \lambda=1 \tag{1.7}
\end{equation*}
$$

at the center of symmetry

$$
\begin{equation*}
f(0, q)=0 \quad \text { for } \lambda=0 \tag{1.8}
\end{equation*}
$$

and the initial conditions

$$
\begin{equation*}
f(\lambda, 0)=f_{0}(\lambda), \quad g(\lambda, 0)=g_{0}(\lambda), \quad h(\lambda, 0)=h_{0}(i) \quad \text { for } q=0 \tag{1.9}
\end{equation*}
$$

where $f_{0}(\lambda), g_{0}(\lambda), h_{0}(\lambda)$ are known functions corresponding to the selfsimilar problem [1,2]. They satisfy the system

$$
\begin{gather*}
\left(f_{0}-\lambda\right) f_{0}^{\prime} \div \frac{2(\gamma-1)}{(\gamma+1)^{2}} \frac{h_{0}^{\prime}}{g_{0}}+\frac{\omega-v}{2} f_{0}=0 \\
\left(f_{0}-\lambda\right) \frac{g_{0}^{\prime}}{g_{0}}+f_{0}^{\prime}+\frac{v-1}{\lambda} f_{0}-\omega=0  \tag{1.10}\\
\left(f_{0}-\lambda\right) \frac{h_{0}^{\prime}}{h_{0}}+\gamma\left(f_{0}^{\prime}+\frac{\nu-1}{\lambda} f_{0}\right)-v=0
\end{gather*}
$$

in addition

$$
\begin{equation*}
f_{0}(1)=\frac{2}{\gamma+1}, \quad g_{0}(1)=h_{0}(1)=1, \quad f_{0}(0)=0 \tag{1.11}
\end{equation*}
$$

Here, as also in all that follows, primes will denote differentiation with respect to $\lambda$. We note that system (1.10) can be obtained from (1.6) by passing to the limit as $q \rightarrow 0$.

For small values of $q$, i.e. for small values of the time, when the explosion is still sufficiently strong, the solution of the problem set down above can be sought in the form

$$
\begin{align*}
& f(\lambda, q)=f_{0}(\lambda) \div q f_{1}(\lambda) \div \ldots \\
& g(\lambda, q)=g_{0}(\lambda)+q g_{1}(\lambda)+\ldots \\
& h(\lambda, q)=h_{0}(\lambda)+q h_{1}(\lambda)+\ldots \tag{1.12}
\end{align*} \quad \frac{R_{2}}{q} \frac{d q}{d R_{2}}=\frac{\nu}{1+A_{1} q+\ldots}
$$

Because a linearized problem is to be solved, then from (1.6), taking into account system (1.10) and neglecting terms of order $q^{2}$ and higher, we obtain the following system of linear equations for determining the functions $f_{1}(\lambda), g_{1}(\lambda), h_{1}(\lambda)$ and the constant $A_{1}$ :

$$
\begin{align*}
& \left(f_{0}-\lambda\right) g_{0} f_{1}^{\prime}+\frac{2(\gamma-1)}{(\gamma+1)^{2}} h_{1}^{\prime}+\left(f_{0}^{\prime}+\frac{\omega+v}{2}\right) g_{0} f_{1}+ \\
& +\left[\left(f_{0}-\lambda\right) f_{0}^{\prime}+\frac{\omega-v}{2} f_{0}\right] g_{1}+\frac{4 \gamma-(\gamma-1)^{2}}{\gamma(\gamma+1)^{2}} h_{0}^{\prime}+\frac{v}{2} f_{0} g_{0} A_{1}=0 \\
& g_{0} f_{1}^{\prime}+\left(f_{0}-\lambda\right) g_{1}^{\prime}+\left(\frac{v-1}{\lambda} g_{0}+g_{0}^{\prime}\right) f_{1}+ \tag{1.13}
\end{align*}
$$

$$
\begin{array}{r}
+\left(\frac{\nu-1}{\lambda} f_{0}+f_{0}^{\prime}+\nu-\omega\right) g_{1}-\frac{2 v}{\gamma-1} g_{0}=0 \\
\Upsilon h_{0} f_{1}^{\prime}+\left(j_{0}-\lambda\right) h_{1}^{\prime}+\left(h_{0}^{\prime}+\gamma \frac{\nu-1}{\lambda} h_{0}\right) f_{1}+\gamma\left(f_{0}^{\prime}+\frac{\nu-1}{\lambda} f_{0}\right) h_{1}- \\
-\nu\left(\frac{\gamma-1}{2 \gamma}-A_{1}\right) h_{0}=0
\end{array}
$$

From conditions (1.7), (1.8), taking into account (1.11) and (1.12), we obtain boundary conditions for the sought for functions $f_{1}(\lambda), g_{1}(\lambda)$ and $h_{1}(\lambda)$ :

$$
\begin{equation*}
f_{1}(1)=-\frac{2}{r+1}, \quad g_{1}(1)=h_{1}(1)=f_{1}(0)=0 \tag{1.14}
\end{equation*}
$$

System (1.13) can be transformed into a form that is more convenient for the investigations which follow. With this aim we shall introduce new unknown functions $F(\lambda), G(\lambda)$ and $H(\lambda)$ which are connected to the functions $f_{1}(\lambda), g_{1}(\lambda)$ and $h_{1}(\lambda)$ by the relations

$$
\begin{equation*}
f_{1}(\lambda)=\left(f_{0}-\lambda\right) F, \quad g_{1}(\lambda)=g_{0} G, \quad h_{1}(\lambda)=h_{0} H \tag{1.15}
\end{equation*}
$$

After the transformation, system (1.3) is written thus:

$$
\begin{align*}
& \left(f_{0}-\lambda\right)^{2} F^{\prime}+\frac{2(\gamma-1)}{(\gamma+1)^{2}} \frac{h_{0}^{\prime}}{g_{0}} H^{\prime}+\left(f_{0}-\lambda\right)\left(2 f_{0}^{\prime}+\frac{\omega+\nu-2}{2}\right) F+  \tag{1.16}\\
& +\frac{2(\gamma-1)}{(\gamma+1)^{2}} \frac{h_{0}^{\prime}}{g_{0}} H+\left[\left(f_{0}-\lambda\right) f_{0}^{\prime}+\frac{\omega-\nu}{2} f_{0}\right] G+\frac{4 \gamma-(\gamma-1)^{2}}{\gamma\left(\gamma^{\prime}+1\right)^{2}} \frac{h_{0}^{\prime}}{g_{0}}+\frac{\nu}{2} f_{0} A_{1}=0 \\
& \quad\left(f_{0}-\lambda\right) F^{\prime}+\left(f_{0}-\lambda\right) G^{\prime}+\left(f_{0}^{\prime}-1\right) F+\left(\frac{\nu-1}{\lambda}+\frac{g_{0}^{\prime}}{g_{0}}\right)\left(f_{0}-\lambda\right) F+ \\
& \quad+\left(\frac{v-1}{\lambda} f_{0}+f_{0}^{\prime}+\nu-\omega\right) G+\left(f_{0}-\lambda\right) \frac{g_{0}^{\prime}}{g_{0}} G-\frac{2 v}{\gamma-1}=0  \tag{1.17}\\
& \quad \gamma\left(f_{0}-\lambda\right) F^{\prime}+\left(f_{0}-\lambda\right) H^{\prime}+\left(f_{0}^{\prime}-1\right) \gamma F+\left(f_{0}-\lambda\right)\left(\frac{h_{0}^{\prime}}{h_{0}}+\gamma \frac{v-1}{\lambda}\right) F+ \\
& \quad+\left(f_{0}-\lambda\right) \frac{h_{0}^{\prime}}{h_{0}} H+\gamma\left(f_{0}^{\prime}+\frac{\nu-1}{\lambda} f_{0}\right) H-v\left(\frac{\gamma-1}{2 \gamma}-A_{1}\right)=0 \tag{1.18}
\end{align*}
$$

From the two latter equations of this system it is possible to obtain a first integral. Let us show this.
2. An integral of system (1.16)-(1.18) and a law of shock wave motion. If the quantities $g_{0}{ }^{\prime} / g_{0}$ and $h_{0}{ }^{\prime} / h_{0}$ be eliminated from equations (1.17) and (1.18), then we will have

$$
\begin{gathered}
\left(f_{0}-\lambda\right)\left(F^{\prime}+G^{\prime}\right)+(\omega-\imath) F+v G-\frac{2 v}{\gamma-1}=0 \\
\left(f_{0}-\lambda\right)\left(\gamma F^{\prime}+H^{\prime}\right)-v(\gamma-1) F+\nu H-v\left(\frac{\gamma-1}{2 \gamma}-A_{1}\right)=0
\end{gathered}
$$

Replacing the independent variable according to the formula

$$
\mu=\int_{\lambda}^{1} \frac{d \lambda}{f_{0}-\lambda}
$$

we obtain

$$
\begin{gathered}
\frac{d}{d \rho^{2}}(F+G)+(\omega-\nu) F+\nu G-\frac{2 \nu}{\gamma-1}=0 \\
\frac{d}{d \mu}(\gamma F+H)-\nu(\gamma-1) F+\nu H-\nu\left(\frac{\gamma-1}{2 \gamma}-A_{1}\right)=0
\end{gathered}
$$

Multiplying now the first equation by $\nu(2 \gamma-1) /(\omega-2 \nu)$ adding to the relation so obtained the second equation and integrating the result, we find

$$
\left[\frac{\nu(2 \gamma-1)}{\omega-2 \nu}+\gamma\right] F+\frac{\nu(2 \sim-1)}{\omega-2 \nu} G+H=C_{1} \exp (-\gamma \mu)+\frac{2 \nu}{\gamma-1} \frac{2 \gamma-1}{\omega-2 \nu}+\frac{\gamma-1}{2 \gamma}-A_{1}
$$

From the boundary conditions

$$
F(1)=\frac{2}{\gamma-1}, \quad G(1)=0, \quad H(1)=0, \quad \mu(1)=0
$$

we find the constant of integration

$$
C_{1}=\frac{(3 \gamma-1)(\gamma+1)}{2 \gamma(\gamma-1)}+A_{1}
$$

If we make use of the adiabatic equation for the self-similar motions

$$
\left(f_{0}-\lambda\right)\left(\frac{h_{0}^{\circ}}{h_{0}}-\gamma \frac{g_{0^{\circ}}}{g_{0}}\right)+\gamma \omega-\nu=0
$$

and if we take into account that $\mu=0, g_{0}=1, h_{0}=1$ for $\lambda=1$, then we obtain

$$
e^{-\mu}=\left(h_{0} / g_{0}^{\gamma}\right)^{\frac{1 \cdot}{\gamma \omega-\nu}}
$$

Thus, a first integral of system (1.16)-(1.18) is found which satisfies the boundary conditions at the shock wave:

$$
\begin{gather*}
{\left[\frac{v(2 \gamma-1)}{\omega-2 v}+\gamma\right] F+\frac{v(2 \gamma-1)}{\omega-2 \nu} G+H=} \\
=\left[\frac{(3 \gamma-1)(\gamma+1)}{2 \gamma(\gamma-1)} A_{1}\right]\left(\frac{h_{0}}{g_{0}^{\gamma}}\right)^{\frac{\nu}{\gamma \omega-v}}+\frac{2 \nu}{\gamma-1} \frac{2 \gamma-1}{\omega-2 \nu}+\frac{\gamma-1}{2 \gamma}-A_{1} \tag{2.1}
\end{gather*}
$$

The existence of an integral analogous to (2.1) was proved by Lidov [6].

With the help of the obtained integral (2.1) the problem reduces to the solution of a system of two linear equations, which for arbitrary values of $\omega$ can be found by numerical integration.

After finding the quantities $A_{1}$ the dependence of $R_{2}(q)$ and $\tau(q)$ can be found, where $R_{2}=r_{2} / r^{0}$ and $\tau=t / t^{0}$. Here $r^{0}$ is a dynamic length, introduced earlier, and $t^{0}$ is a dynamic time defined by the formula

$$
t^{\circ}=E_{0}{ }^{1 / \nu} \boldsymbol{A}^{1 / 2} \quad r^{0-\omega / 2} p_{1}^{-(v+2) / 2 v}
$$

It is known [1] that for the self-similar solution we have

$$
\begin{equation*}
r_{2}(t)=\left(\frac{E_{0}}{\alpha A}\right)^{8 / 2} t^{\delta}, \quad c^{2}=\delta^{2} r_{2}^{2} t^{-2}, \quad \delta=\frac{2}{v+2-\omega} \tag{2.2}
\end{equation*}
$$

where $a(y, \omega)$ is a known quantity.
If equations (2.2) be transformed to the introduced non-dimensional parameters, then the dependence of $R_{2}(q)$ and $\tau(q)$ can be found in the self-similar problem:

$$
\begin{equation*}
R_{2}{ }^{v}(q)=\frac{\delta^{2}}{\gamma \alpha} q, \quad \tau(q)=\left(\frac{\delta^{2}}{\gamma \alpha}\right)^{\frac{2-v \delta}{2 \cdot \delta}}\left(v \gamma^{1 / 2}\right)^{-1} q^{\frac{1}{\nu \delta}} \tag{2.3}
\end{equation*}
$$

For the linearized problem of (1.12), by integrating and taking into account (2.3), we obtain

$$
\begin{equation*}
R_{2}^{\nu}(q)-\frac{\delta^{2}}{\gamma \alpha} q \exp \left(A_{1} q\right) \tag{2.4}
\end{equation*}
$$

Let us find $\tau(q)$. Using the definitions of $q, \tau, R_{2}$, it is easy to show that

$$
\frac{d R_{2}}{d \tau}=\left(\frac{\gamma R_{2}^{\omega}}{q}\right)^{1 / 2}
$$

Since $d t / d q=\left(d R_{2} / d q\right)\left(d r / d R_{2}\right)$, taking into account (2.4) we find

$$
\frac{d \tau}{d q}=\left(\frac{\delta^{2}}{\gamma \alpha}\right)^{\frac{2-\omega}{2 v}}\left(\nu \gamma^{2 / 2}\right)^{-1}\left(1+A_{1} q\right) q^{\frac{2-v-\omega}{2 v}} \exp \left(\frac{2-\omega}{2 v} A_{1} q\right)
$$

For small values of $q$ it is possible to write

$$
\left(1+A_{1} q\right) \exp \left(\frac{2-\omega}{2 \nu} A_{1} q\right)=1+\frac{2 \nu+2-\omega}{2 v} A_{1} q
$$

Thus, for the determination of $r(q)$ we obtain the differential equation

$$
\frac{d \tau}{d q}=\left(\frac{\delta^{2}}{\gamma \alpha}\right)^{\frac{2-\omega}{2 v}}\left(\nu \gamma^{1 / 2}\right)^{-1}\left[1+\frac{2 v+2-\omega}{2 v} A_{1} q\right]_{q^{\frac{2-v-\omega}{2 v}}}^{\frac{2 v}{2}}
$$

Integrating this equation and determining the constant of integration from the condition $r(0)=0$, we find the required dependence

$$
\begin{equation*}
\tau(q)=\left(\frac{\delta^{2}}{\gamma \alpha}\right)^{\frac{2-v \delta}{2 v \delta}}\left(v \gamma^{1 / s}\right)^{-1} q^{\frac{1}{v \delta}}\left[v \delta+\frac{v \delta+2}{2(v \delta+1)} A_{1} q\right] \tag{2.5}
\end{equation*}
$$

Relations (2.4) and (2.5) give in parametric form the law of the shock wave motion, i.e. the dependence of $R_{2}(r)$.

Using (2.4) and (2.5) and the conditions at the shock wave (1.2), it is possible to determine the dependence of all characteristics of the shock wave front from its radius and the time.
3. Exact solution of the problem for $\omega=\omega_{1}$. It was shown earlier that the solution of the linearized problem of an explosion in a medium with variable density can always be obtained by numerical integration of system (1.16)-(1.18). However, in the case for which

$$
\omega=\omega_{1}=\frac{3 v-2+\gamma(2-v)}{\gamma+1}
$$

the solution of this problem can be given in the form of closed formulas. This can be explained since for this value of $\omega$ the self-similar solution has the simple form

$$
\begin{equation*}
f_{0}(\lambda)=\frac{2}{\gamma+1} \lambda, \quad g_{0}(\lambda)=\lambda^{\nu-2}, \quad h_{0}(\lambda)=\lambda^{\nu} \tag{3.1}
\end{equation*}
$$

Substituting $f_{0}(\lambda), g_{0}(\lambda), h_{0}(\lambda)$ from (3.1) into the coefficients of equations (1.16)-(1.18), we find a system of three ordinary non-homogeneous equations with coefficients which depend on the parameters $\gamma$ and $\nu$.

$$
\begin{align*}
& \lambda F^{\prime}+\frac{2}{\gamma-1} \lambda H^{\prime}-\frac{2(\nu+1)}{\gamma-1} F-\frac{2 \nu}{\gamma-1} G+\frac{2 \nu}{\gamma-1} H+ \\
& \quad+\frac{\nu}{(\gamma-1)^{2}}\left[\frac{4 \gamma-(\gamma-1)^{2}}{\gamma}+A_{1}(\gamma+1)\right]=0 \\
& \lambda F^{\prime}+\lambda G^{\prime}+2(\gamma-1) F+\frac{2 v(\gamma+1)}{(\gamma-1)^{2}}-\frac{\nu(\gamma+1)}{\gamma-1} G=0  \tag{3.2}\\
& \lambda \gamma F^{\prime}+\lambda H^{\prime}+\nu(\gamma+1) F-\frac{\nu(\gamma+1)}{\gamma-1} H+\frac{\gamma+1}{\gamma-1} \nu\left(\frac{\gamma-1}{2 \gamma}-A_{1}\right)=0
\end{align*}
$$

This system is a system of equations with constant coefficients, if we take $l=\ln \lambda$ as the independent variable.

For the solution of system (3.2), it would be possible to use integral (2.1) which for $\omega=\omega_{1}$ has the form:

$$
\begin{gather*}
{\left[\frac{\nu(2 \gamma-1)}{\omega_{1}-2 \nu}+\gamma\right] F+\frac{\nu(2 \gamma-1)}{\omega_{1}-2 \nu} G+H=} \\
=\left[\frac{(3 \gamma-1)(\gamma+1)}{2 \gamma(\gamma-1)}+A_{1}\right] \lambda^{\frac{\nu(\gamma+1)}{\gamma-1}}+\frac{2 v}{\gamma-1} \frac{2 \gamma-1}{\omega_{1}-2 v}+\frac{\gamma-1}{2 \gamma}-A_{\perp} \tag{3.3}
\end{gather*}
$$

But because the complete system (3.2) is easily integrated, one need not use the integral (3.3). However, it can be useful in computing the
dependence of the required functions on $\lambda$ and also as a check on the calculation.

We shall find the general solution of the homogeneous system of equations (3.2). The auxiliary equation of the system (3.2) is written thus:

$$
\left|\begin{array}{ccc}
n-\frac{2(v+1)}{\gamma-1} & -\frac{2 v}{\gamma-1} & \frac{2}{\gamma-1}(n+v)  \tag{3.4}\\
n+2(v-1) & n-\frac{v(\gamma+1)}{\gamma-1} & 0 \\
\gamma n+v(\gamma+1) & 0 & n-\frac{v(\gamma+1)}{\gamma-1}
\end{array}\right|=0
$$

To each root $n_{i}$ of equation (3.4) there corresponds a solution $\exp \left(n_{i} l\right)=\lambda^{n_{i}}$ of the homogeneous system.

The first root of equation (3.4) is equal to

$$
\begin{equation*}
n_{1}=\nu(\gamma+1) /(\gamma-1) \tag{3.5}
\end{equation*}
$$

The second and third roots of equation (3.4) satisfy the following quadratic equation:

$$
n^{2}+\frac{5 v \gamma+3 v+\dot{2}}{\gamma+1} n-\frac{2 v(v \gamma+1)(3-\gamma)}{\gamma^{2}-1}=0
$$

and are equal to

$$
\begin{equation*}
n_{2}=-\frac{b_{1}}{2}+\sqrt{\frac{b_{1}^{2}}{4}+b_{2}}, \quad n_{3}=-\frac{b_{1}}{2}-\sqrt{\frac{b_{1}^{2}}{4}+b_{2}} \tag{3.6}
\end{equation*}
$$

where

$$
b_{1}=\frac{5 v \gamma+3 v+2}{\gamma+1}, \quad b_{2}=\frac{(2 v(v \gamma+1)(3-\gamma)}{\gamma^{2}-1}
$$

From (3.6) it is seen that $n_{2}=0$ for $\gamma=3$. The dependence of the roots $n_{1}, n_{2}$ and $n_{3}$ on $\nu$ and $\gamma$ are given in the Table.

The general solution of a homogeneous system corresponding to system (3.2) can be represented in the form

$$
\begin{gather*}
F(\lambda)=c_{2} \lambda^{n_{2}}+c_{3} \lambda^{n_{3}} \\
G(\lambda)=c_{1} \lambda^{n_{1}}+\frac{\left(n_{2}+2 \nu-2\right)(\gamma-1)}{(\gamma+1) \nu-(\gamma-1) n_{2}} c_{2} \lambda^{n_{2}}+\frac{\left(n_{3}+2 \nu-2\right)(\gamma-1)}{(\gamma+1) \nu-(\gamma-1) n_{3}} c_{3} \lambda^{n} \quad  \tag{3.7}\\
H(\lambda)=\frac{\gamma-1}{2 \gamma} f_{1} \lambda^{n_{1}}+\frac{\left(\gamma n_{2}+\nu \gamma-1-\nu\right)(\gamma-1)}{(\gamma+1) \nu-(\gamma-1) n_{2}} c_{2} \lambda^{n_{2}}+\frac{\left(\gamma n_{3}+\nu \gamma+\nu\right)(\gamma-1)}{(\gamma+1) \nu-(\gamma-1) n_{9}} c_{3} \lambda^{n_{3}}
\end{gather*}
$$

The particular solution of the non-homogeneous system of equations (3.2) will thus be

$$
\begin{equation*}
F=\alpha_{1}, \quad G=\alpha_{2}, \quad H=\alpha_{3} \tag{3.8}
\end{equation*}
$$

where $a_{1}, a_{2}, a_{3}$ are expressed in terms of $A_{1}, \nu$ and $\gamma$ according to the formulas

$$
\begin{gather*}
\alpha_{1}=\frac{v(\gamma+1)}{2(\gamma-1)(v \gamma+1)} A_{1}, \quad \alpha_{2}=\frac{v-1}{v \gamma+1} A_{1}+\frac{2}{\gamma-1}  \tag{3.9}\\
\alpha_{3}=\frac{1}{2}\left[\frac{\gamma-1}{\gamma}-\frac{v \gamma-v+2}{v \gamma+1} A_{1}\right]
\end{gather*}
$$

Using (3.7) and (3.8) we find the solution of system (1.13) for $\omega=\omega_{1}$ in the following form

$$
\begin{align*}
& f_{1}(\lambda)=\frac{1-\gamma}{\gamma+1} \lambda\left[\alpha_{1}+c_{2} \lambda^{n_{2}}+c_{3} \lambda^{n_{3}}\right] \\
& g_{1}(\lambda)=\lambda^{v-2}[ \alpha_{2}+c_{1} \lambda^{n_{1}}+\frac{\left(n_{2}+2 v-2\right)(\gamma-1)}{v(\gamma+1)-(\gamma-1) n_{2}} c_{2} \lambda^{n_{2}}+  \tag{3.10}\\
&\left.+\frac{\left(n_{3}+2 v-2\right)(\gamma-1)}{v(\gamma+1)-(\gamma-1) n_{3}} c_{3} \lambda^{n_{1}}\right] \\
& h_{1}(\lambda)=\lambda^{\nu}\left[\alpha_{3}+\frac{\gamma-1}{2 \gamma} c_{1} \lambda^{n_{1}}+\frac{\left(\gamma n_{2}+v \gamma+v\right)(\gamma-1)}{v(\gamma+1)-(\gamma-1) n_{2}} c_{2} \lambda^{n_{2}}+\right. \\
&\left.+\frac{\left(\gamma n_{3}+v \gamma+v\right)(\gamma-1)}{v(\gamma+1)-(\gamma-1) n_{3}} c_{3} \lambda^{n_{2}}\right]
\end{align*}
$$

We shall determine the constants $c, c_{2}, c_{3}$ and $A_{1}$ so that the functions $f_{1}(\lambda), g_{1}(\lambda), h_{1}(\lambda)$ satisfy the boundary conditions (1.14).

Thus, from the latter equation of (1.14) (the velocity at the center must be equal to zero), taking into consideration that $n_{3}$ for arbitrary $\gamma$ is a negative quantity whose modulus is greater than unity, we obtain $c_{3}=0$; from the other equations of (1.14) we obtain a system of nonhomogeneous linear equations with coefficients which depend on $\gamma$ and $\nu$.

$$
\begin{gather*}
\alpha_{1}+c_{2}+\frac{2}{\gamma-1}=0  \tag{3.11}\\
\alpha_{2}+c_{1}+\frac{\left(n_{2}+2 v-2\right)(\gamma-1)}{v(\gamma+1)-(\gamma-1) \cdot n_{2}} c_{2}=0  \tag{3.12}\\
\alpha_{3}+\frac{\gamma-1}{2 \gamma} c_{1}+\frac{\left(\gamma n_{2}+v \gamma+v\right)(\gamma-1)}{v(\gamma+1)-(\gamma-1) n_{2}} c_{2}=0 \tag{3.13}
\end{gather*}
$$

4. Results of calculations according to formulas of the exact solution. If $c_{1}$ from (3.12) is substituted into equation (3.13) and then equations (3.11) and (3.9) are used, we obtain a relation for determining $A_{1}$. The quantity $A_{1}$ will depend on $\gamma$ and $\nu$.

After determining $A_{1}$ from the system of equations (3.9), (3.11), (3.12), we find the dependence on $\gamma$ and $\nu$ of the quantities $a_{1}, a_{2}, a_{3}, c_{1}, c_{2}$.

The calculation formulas for finding the indicated quantities have the form

$$
\begin{aligned}
& A_{1}=\frac{B_{6}}{B_{5}}, \quad \alpha_{1}=\frac{1}{\gamma-1} B_{3} A_{1}, \quad \alpha_{2}=\frac{\nu-1}{v \gamma+1} A_{1}+\frac{2}{\gamma-1} \\
& x_{3}=\frac{\gamma-1}{2 \gamma}-B_{4} A_{1}, \quad c_{2}=\frac{2}{\gamma-1}-\alpha_{1}, \quad c_{1}=-\left[\alpha_{2}+(\gamma-1) B_{1} c_{2}\right]
\end{aligned}
$$

where

$$
\begin{aligned}
& B_{1}=\frac{n_{2}+2 v-2}{v(\gamma+1)-(\gamma-1) n_{2}}, \quad B_{2}=\frac{\gamma n_{2}+\nu \gamma+v}{\nu(\gamma+1)-(\gamma-1) n_{2}} \\
& B_{3}=\frac{v(\gamma+1)}{2(v \gamma+1)}, \quad B_{4}=\frac{v \gamma-v+2}{2(v \gamma+1)} \\
& B_{5}=\frac{\gamma-1}{2 \gamma}\left(B_{1} B_{3}-\frac{\nu-1}{\nu \gamma-1}\right)-B_{2} B_{3}-B_{4}, \quad B_{6}=\frac{3-\gamma}{2 \gamma}+\frac{\tau-1}{\gamma} B_{1}-2 B_{2}
\end{aligned}
$$

TABLE 1.

| $\checkmark$ | $\gamma$ | $\omega_{1}$ | $n_{1}$ | $n_{2}$ | $-n$, | $\alpha_{1}$ | $\alpha_{2}$ | $-x_{3}$ | $-\mathrm{C}_{1}$ | $c_{2}$ | $A_{1}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 3 | 1.2 | 2.6364 | \|33.000| | 5.9131 | 19.095 | 7.0942 | 10.860 | 0.4756 | $11.9 \% 3$ | 2.9058 | 1.9778 |
|  | 1.4 | 2.3333 | 18.000 | 3.1540 | 16.487 | 3.2461 | 5.72130 | 0.4342 | 6.5665 | 1.7540 | 1.8755 |
|  | $5 / 3$ | 2.0000 | 12.000 | 1.7684 | 15.268 | 1.8211 | 3.6070 | 0.4070 | 4.2717 | 1.1789 | 1.8211 |
|  | 3.0 | 1.0000 | 6.000 | 0.0000 | 14.000 | 0.5333 | 1.3556 | 0.3778 | 1.6667 | 0.4667 | 1.7778 |
|  | 7.0 | 0.0000 | 4.000 | $-0.8031$ | 13.697 | 0.1635 | 0.4968 | 0.3887 | 0.6098 | 0.1699 | 1.7979 |
| 2 | 1.2 | 1.8182 | 22.000 | 4.1894 | 1.3 .280 | 6.6951 | 10.609 | 0.6470 | 13.228 | 3.3049 | 2.0694 |
|  | 1.4 | 1.6667 | 12.000 | 2.2240 | 11.391 | 3.1520 | 5.52530 | 0.5926 | 6.3238 | 1.8481 | 1.9962 |
|  | $5 / 3$ | 1.5000 | 8.000 | 1.2394 | 10.489 | 1.7715 | 3.4429 | 0.5381 | 4.0315 | 1.2285 | 1.9191 |
|  | 3.0 | 1.0000 | 4.000 | 0.0000 | 9.500 | O. 5238 | 1.2619 | 0.4524 | 1.5000 | 0.4762 | 1.8333 |
|  | 7.0 | 0.5000 | 2.667 | $-0.5431$ | 9.207 | 0.1621 | 0.4549 | 0.4225 | 0.5326 | 0.1712 | 1.8237 |

The results of the calculations are presented in the table.
The formulas (3.10) taking into account that $c_{3}=0$ are written thus:

$$
\begin{gather*}
f_{1}(\lambda)=\frac{1-\tau}{\gamma+1} \lambda\left(\alpha_{1}+c_{2} \lambda^{n_{2}}\right) \\
g_{1}(\lambda)=\lambda^{\nu-2}\left[\alpha_{2}+c_{1} \lambda^{n_{1}}+(\gamma-1) B_{1} c_{2} \lambda^{n_{2}}\right]  \tag{4.1}\\
h_{1}(\lambda)=\lambda^{\nu}\left[\alpha_{3}+\frac{\gamma-1}{2 \gamma} c_{1} \lambda^{n_{1}}+(\gamma-1) B_{2} c_{3} \lambda^{n_{2}}\right]
\end{gather*}
$$

From the relation (0.3) in the case of spherical waves ( $\nu=3$ ), cylindrical waves $(\nu=2)$ and plane waves $(\nu=1)$ we will have respectively

$$
\omega_{1}=\frac{7-\gamma}{\gamma+1}, \quad \omega_{1}=\frac{4}{\gamma+1}, \quad \omega_{1}=1
$$

The formula for finding the mass of matter enclosed in some findte region containing the origin of the coordinate system

$$
m=\sigma_{v} \int_{0}^{r} p r^{\nu-1} d r
$$

shows that the initial mass will be of finite magnitude only in the spherical and cylindrical cases. Therefore, the calculations were carried out for these two cases.

The integral (3.3) in the case of spherical symmetry takes the form

$$
\begin{gathered}
\frac{(\gamma-3)(\gamma-1)}{\overline{1}-1} F-\frac{3(2 \gamma-1)(\gamma+1)}{7 \gamma-1} G+H= \\
=\left[\frac{(3 \gamma-1)(\gamma+1)}{2 \gamma(\gamma-1)}+A_{1}\right]^{\frac{3(\gamma+1)}{\gamma-1}}-\frac{6}{\gamma-1} \frac{(2 \gamma-1)(\gamma+1)}{7 \gamma-1}+\frac{\gamma-1}{2 \gamma}-A_{1}
\end{gathered}
$$

and in the cylindrical case

$$
\begin{gathered}
-\frac{\gamma-1}{2 \gamma} F-\frac{(2 \gamma-1)(\gamma+1)}{2 \gamma} G+H=\left[\frac{(3 \gamma-1)(\gamma+1)}{2 \gamma(\gamma-1)}+A_{1}\right] \lambda^{\frac{2(\gamma+1)}{\gamma-1}}+ \\
+\frac{\gamma-1}{2 \gamma}-\frac{(2 \gamma-1)(\gamma+1)}{\gamma(\gamma-1)}-A_{1}
\end{gathered}
$$

The dependence of $R_{2}(q)$ and $t(q)$ is found from the relations (2.4) and (2.5) in which it is necessary to set $\delta=\delta_{1}$, where

$$
\delta_{1}=\delta\left(\omega_{1}\right)=\frac{\gamma+1}{v \gamma-v+2}
$$

Using the calculated constants and formulas (4.1), the dependence of $f_{1}(\lambda), g_{1}(\lambda)$ and $h_{1}(\lambda)$ (Figs. $1,2,3$ respectively) were constructed for various values of $\gamma$ and $\nu$. With the help of the functions $f_{1}(\lambda), g_{1}(\lambda)$ and $h_{1}(\lambda)$, and knowing the constant $A_{1}$, it is possible to calculate the characteristics of the motion for small values of $q$.

$\boldsymbol{a}$

b

Fig. 1.
In Figs. $4 a$ and $4 b$ the distribution of the non-dimensional pressure $h=h_{0}+q h_{1}$ is given in the spherical case for $\gamma=1.4$ and $\gamma=7$ and in
the cylindrical case for $y=1.4$ for the values $q=0$ and $q=0.2$.


Fig. 2.


Fig. 3.
These graphs show the influence of back pressure on the development of an explosion in the initial stage.

If the values for $\omega_{1}$ found for $\nu=3$ and $\nu=2$ are equated, then we obtain $\gamma=3$. In this case $\omega_{1}=1$. A comparison of the values of the nondimensional pressure for $\gamma=3, \omega_{1}=1$ and various $\nu(\nu=3, \nu=2)$ is
given in Fig. 5.


Fig. 4.

We note in conclusion that the exact solution for $\omega=\omega_{1}$ in the particular case $\nu=3$ obtained in the present paper supplements and makes more precise the results of [7] (see also [8]).


Fig. 5.
The author sincerely thanks N. S. Mel'nikov for discussing certain questions which are examined in the present paper.

## BIBLIOGRAPHY

1. Sedov, L. I., Metody podobiia i ramernosti vekhanike (Methods of Siailarity and Dimensional Analysis in Mechanics). 4th Ed. Moscow, 1957.
2. Korobeinikov, V.P. and Riazanov, E.V., Predstavlenie reshenifa zadachi o tochechnom varyve $v$ gaze $v$ osobykh sluchaiakh (Presentation of a solution of the problem of a point explosion in a gas in special cases). PMM Vol. 23, No. 2, 1959.
3. Okhotsimskii, D.E., Kondrasheva, 1.L., Vlasova, Z. P. and Kazakova, R.K., Raschet tochecknogo vzryva s uchetom protivodavleniia (Calculation of a point explosion including back pressure). Tr. Maten. inuta Akad. Nauk SSSR, Vol. 50, 1957.
4. Sakurai, A., On propagation and structure of the blast wave. J. Phys. Soc. Japan, Vol. 8, No. 5, 1953; Vol. 9, No. 2, 1954.
5. Brushlinskii, D. N. and Solomakhova, T.S., Issledovanie zadachi o cil'nom varyve s uchetom protivodavleniia (Investigation of the problem of a strong explosion including back pressure). Sb. st. No. 19, No. 7, Theoretical Hydrodynamics, edited by L. I. Sedov, Oborongiz, 1956.
6. Lidov, M.L., $K$ teorii linearizirovannykh reshenii okolo odnomernykh avtomodel'nykh dvizhenii gaza (On the theory of the linearized solution of the one-dimensional self-modeling motion of a gas). Dokl. Akad. Nauk SSSh, Vol. 102, No. 6, 1955.
7. Korobeinikov, V.P. and Mel'nikova, N. S., 0 tochnykh resheniiakh linearizirovannoi zadachi o tochechnom vzryve s protivodavleniem (On exact solutions of a linearized problem of a point explosion with back pressure). Dokl. Akad. Nauk SSSR, Vol. 116, No. 2, 1957.
8. Korobeinikov, V.P. and Mel'nikova, N.S., Pis'mo v redaktsilu (Letter to the editor). Dokl. Akad. Nauk SSSR, Vol. 126, No. 1, 1959.
